

esas

May 4, 2016

Abstract

This document describes the Extended Source Analysis Software (XMM-ESAS) package for the analysis of EPIC MOS and PN observations. Originally ESAS was a stand-alone package, which relied on SAS, was comprised of FORTRAN 77 routines and Perl scripts. It is now incorporated into SAS. ESAS includes routines which create source and model particle background spectra and exposure-corrected, background-subtracted (particle, soft proton, and solar wind charge exchange) images. The spectra and images are produced for user-defined regions within an observation field of view. The output files are in standard FITS format. Software for mosaicking multiple observations of not necessarily co-aligned observations is included in this package. Note: This documentation is meant to complement the document COOKBOOK FOR ANALYSIS PROCEDURES FOR XMM-NEWTON EPIC OBSERVATIONS OF EXTENDED OBJECTS AND THE DIFFUSE BACKGROUND[3].

1 Instruments/Modes

Instrument	Mode
EPIC MOS	IMAGING
EPIC PN	IMAGING

2 Use

pipeline processing	no	
interactive analysis	yes	

3 Description

This package consists of a collection of XMM-Newton Extended Source Analysis Software (XMM-ESAS) tasks originally developed by Snowden, et al. and released as a stand-alone package in 2008. This package follows the methods outlined in Snowden et al. (2008)[1] for the analysis of extended objects and the diffuse background using XMM-Newton EPIC MOS and pn observations.

Two separate features are incorporated into ESAS; the capability of creating model quiescent particle background spectra (Kuntz & Snowden 2008)[2] for user defined regions of the detectors and the capability

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of creating background subtracted and exposure corrected images. Also included in the XMM-ESAS package is software to mosaic multiple and not necessarily coaligned observations of regions on the sky.

SAS/ESAS Task Descriptions:

FORTRAN routines:

• adapt - FORTRAN

adapt is an adaptive filtering routine used to create smoothed background subtracted and exposure corrected images for individual exposures or exposures from a single observation combined by the task *comb*. For each unmasked pixel, the program will average neighboring pixels within a circle of increasing radius until a selected number of counts is reached. The original pixel is then given the average surface brightness for the pixels within the circle. Images can also be binned before smoothing.

• $adapt_merge - FORTRAN$

adapt_merge adaptively smooths background subtracted and exposure corrected mosaicked images using the output of merge_comp_xmm. For each unmasked pixel, the program will average neighboring pixels within a circle of increasing radius until a selected number of counts from the count image is reached. The original pixel is then given the average surface brightness for the pixels within the circle. The images can also be binned before smoothing.

• bin_image - FORTRAN

bin_image produces binned count rate and count-rate uncertainty images of single observations. It can use either individual exposures or the output of the program comb which can merge all of the exposures associated with a single ObsID. For each unmasked and binned pixel, the program will determine the average count rate and the count rate uncertainty. The assumption is that the uncertainty is dominated by the counting statistics and the the systematics of the background modeling. Integer binning, including by 1 for no binning, is supported.

• bin_image_merge - FORTRAN

bin_image_merge bins mosaicked images combined by the task merge_comp_xmm into count rate and count rate uncertainty images. For each unmasked and binned pixel, the program will determine the average count rate and the count rate uncertainty. The assumption is that the uncertainty is dominated by the counting statistics and the the systematics of the background modeling. Integer binning, including by 1 for no binning, is supported.

• clean - PERL

clean deletes a number of intermediate and unneeded files after the processing is complete.

• cheese - PERL

cheese does source detection and creates cheese masks for point-source masking during image processing. *cheese* allows the user to set a flux threshold over the field for the removal of the source contributions to the spectra as well as the images.

• cheese-bands - PERL

cheese-bands does source detection and creates cheese masks for point-source masking during image processing in three bands: soft, hard, and combined. cheese-bands allows the user to set a flux threshold over the field for the removal of the source contributions to the spectra as well as the images.



• comb - FORTRAN

comb combines co-aligned event, exposure, QPB, SP, and SWCX background images from different exposures and different instruments from the same ObsID.

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• $conv_reg$ - FORTRAN

conv_reg converts region information in celestial coordinates into region information in detector coordinates. It will do so for region fits files, ascii lists, or individual regions. **Development in progress.**

• conv-region — PERL

conv-region converts a region file in celestial coordinates into regions files in detector coordinates for all active detectors in all observations in a list of ObsIDs. **Development in progress.**

• esas-mosaic - PERL

 $esas_mosaic$ is the driving task for $esprep_mosaic$ that will process all of the exposures for an individual subpointing for all of the subpointings for an observation in mosaic mode.

• esprep_mosaic - FORTRAN

 $esprep_mosaic$ takes the output event files for an individual exposure in an individual sub-pointing of a mosaic mode observation and resets the X,Y coordinates and pointing-direction keywords to make them compatible with the ESAS package. It is called by the task esas-mosaic

• esprep_mosaic_time - FORTRAN

esprep_mosaic_times extracts the position angle, right ascension, and declination from individual subpointings of mosaic mode observations for use in esprep_mosaic.

• $make_mask - FORTRAN$

make_mask is called by the task cheese to produce a cheese mask.

• $make_mask_merge$ - FORTRAN

make_mask_merge produces cheese masks in the total band (cheese) or soft, hard, and combined bands (cheese-bands) using a merged source list produced by the task merge-source-list from the maximum likelihood source lists produced by cheese or cheese-bands.

• $merge_comp_xmm$ - FORTRAN

merge_comp_xmm creates mosaicked count, exposure, QPB, SP, and SWCX background images for multiple observations and exposures.

• $merge_source_list$ - FORTRAN

merge_source_list creates a master list of sources detected from a list of observations which have been processed to the point of a creating an emldetect source lists produced by *cheese* or *cheese-bands*. Sources are compared and those within 2" of each other have the one with the lower flux thrown out.

• mos_back - FORTRAN

mos_back takes the output of the perl script mos-spectra and creates quiescent particle background (QPB) spectra and images in detector coordinates, as selected, for EPIC MOS data.

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 \bullet mos-filter - PERL

mos-filter filters and cleans the event files of SP contamination using the SAS task esp-filt. Along with a filtered event file it produces a QDP plot file showing the observation light curves and indicates the accepted time intervals.

 \bullet mos-spectra - PERL

mos-spectra processes the filtered event files from the task mos-filter to produce a set of intermediate files for the production of QPB background spectra and images. mos-spectra also produces source spectra and the appropriate Redistribution Matrix Files, RMFs, and Ancillary Region Files, ARFs, for spectral analysis.



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• pn_back - FORTRAN

 pn_back takes the output of $pn_spectra$ and creates QPB spectra and images in detector coordinates, as selected, for EPIC pn data.

• pn-filter - PERL

pn-filter is the initial processing script for pn data and it performs the same tasks as mos-filter for MOS data. Each exposure is processed in both normal and out-of-time (OOT) modes.

• pn-spectra - PERL

pn-spectra provides the same functions for pn data as mos-spectra provides for the MOS. However, it also creates spectra and images from the out-of-time processing.

• point_source - FORTRAN

point_source calculates the appropriate Xspec normalization for the extragalactic background given the user-selected point-source exclusion threshold.

• proton - FORTRAN

proton produces images in detector coordinates of the model residual soft proton (SP) contamination. To do so, it uses the fitted values of the SP component from Xspec and standard detector maps.

• proton_scale - FORTRAN

proton_scale extracts the appropriate scale factors for the solid angle and relative SP contributions to include in spectral fitting.

• rot_det_sky - FORTRAN

 rot_det_sky is called by the $rot_im_det_sky$ task. It does the heavy lifting of rotating the QPB images output from mos_back and pn_back , the SP images output from proton, and the SWCX output from swcx which are in detector coordinates, into images in sky coordinates.

• rot-im-det-sky — PERL

rot-im-det-sky rotates both the model QPB, SP, and SWCX background images from detector coordinates to sky coordinates.

• $sp_partial$ - FORTRAN

 $sp_partial$ scales the SP spectral fit results from a limited region of the detector to the full field of view. This allows the user to fit for the SP contamination in regions of lower surface brightness (e.g., an outer annulus for a cluster of galaxies) which can significantly improve the accuracy of the fit.

• swcx - FORTRAN

swcx produces images in detector coordinates of the model residual SWCX contamination. To do so, it uses the fitted values of the SWCX component from Xspec and standard detector maps.



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4 Parameters

This section documents	the parameters recogn	ized by this task	(if any).
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D I	3.6 1	TD.	D C 1	0
Parameter	Mand	Type	Default	Constraints

1. Task adapt parameters:

smoothingcounts	yes	int	50	

The number of counts to accumulate for the smoothing

thresholdmasking	yes	real	0.02	

The scale factor for excluding regions from the smoothing based on a mask image. In the default mode the average exposure is calculated and then any pixel with exposure less than fraction*average value is excluded.

detector	yes	$_{ m int}$	0	0—1

Detector, 1 for a specific instrument and exposure, 0 for the combined image (i.e., the output of comb).

elow	yes	int	400	

The low energy for the band in eV

ehigh	yes	int	1250	

The high energy for the band in eV

binning	yes	int	1	
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Binning control with 1 for no binning, and integers greater than 1 for binning that number of pixels in each dimension.

withpartcontrol	yes	bool	yes	
D 1 1 1 1 1	1 11 11	1	.1 11 . 11	1 1 .

Particle background control, "yes" to subtract the model particle background image.

withsoftcontrol	yes	bool	no	
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Soft proton background control, "yes" to subtract the soft proton background image.

withswcxcontrol	yes	bool	no	
Solar wind charge exchan	ge backgro	and contro	ol. "ves" to subtract the	he SWCX background

Solar wind charge exchange background control, "yes" to subtract the SWCX background image.

	1			
withmaskcontrol	yes	bool	no	

Control for including an additional masking image.

maskfile	yes	dataset	

The file name for an image to provide additional masking if desired. If left blank then there will be no additional masking. The mask images must be the same size and projection as the other images.

prefix	yes	string	1S001	
Prefix defining the exposi	S003 means PN S003			

exposure, while 1S002 and 2S003 mean MOS1 S002 and MOS2 S003 exposures, respectively.

clobber	no	boolean	yes	T/F
C1 1 1				

Clobber existing files?



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2. Task adapt_merge parameters:

smoothingcounts yes int 100	
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The number of counts to accumulate for the smoothing

thresholdmasking yes real 0.02

The scale factor for excluding regions from the smoothing based on a mask image. In the default mode the average exposure is calculated and then any pixel with exposure less than fraction*average value is excluded.

elowlist yes	int	400 750	
--------------	-----	---------	--

Low energy for successive bands in eV

ehighlist	yes	int	750 1250	
-----------	-----	-----	----------	--

High energy for successive bands in eV

binning yes int 1

Binning control, number of pixels (in both dimensions) to be binned.

withpartcontrol yes bool yes

Particle background control, "yes" to subtract the model particle background image.

withsoftcontrol yes bool no

Soft proton background control, "yes" to subtract the soft proton background image.

withswcxcontrol yes bool no

Solar wind charge exchange background control, "yes" to subtract the SWCX background image.

withoffsetbkgcontrol yes bool yes

Offset background control, "yes" to subtract the offset background image. This is a feature currently under development and is not yet functional.

withmaskcontrol yes bool yes

Mask control, "yes" for using a mask image (pixel with 1 in image will be included, pixel with 0 will be excluded).

mask yes dataset mask.fit

Mask image file name.

fill yes int 1

Number of passes to fill in empty pixels. If a zero pixel has three or more non-zero neighbors, the pixel will be the average value of those neighbors.

clobber	no	boolean	yes	T/F

Clobber existing files?

3. Task bin_image parameters:

${ m threshold masking}$	yes	real	0.02	

The scale factor for excluding regions from the smoothing based on a mask image. In the default mode the average exposure is calculated and then any pixel with exposure less than fraction*average value is excluded.



detector yes int Detector selection, 0: combined exposures, 1: MOS, 2: PN. 1S001yes string Prefix defining the exposure used, with the esas nomenclature, eg. S003 means PN S003 exposure, while 1S002 and 2S003 mean MOS1 S002 and MOS2 S003 exposures, respectively. elow int 400 yes Low energy for band in eV 1250 ehigh intHigh energy for band in eV binning int yes Binning control with 1 for no binning, other integers for binning. withpartcontrol bool yes yes Particle background control, "yes" to subtract the model particle background image. withsoftcontrol bool yes no to subtract the soft proton background image. Soft proton background control, "yes" withswcxcontrol yes bool no Solar wind charge exchange background control, "yes" to subtract the SWCX background image. withmaskcontrol bool Solar wind charge exchange background control, "yes" to subtract the SWCX background image. mask dataset mask.fit yes Mask image file name (defaults to using exposure mask). T/Fclobber boolean no yes Clobber existing files? 4. Task bin_image_merge parameters: thresholdmasking 0.02 yes real The scale factor for excluding regions from the smoothing based on a mask image. In the default mode the average exposure is calculated and then any pixel with exposure less than fraction*average value is excluded. elowlist 350 800 yes int Low energy for successive bands in eV ehighlist 800 1300 yes High energy for successive bands in eV

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Particle background control, "yes" to subtract the model particle background image.

yes

bool

int Binning control with 1 for no binning, 2,4,8,16,32 for binning by 2, 4,8,16,32.

yes

binning

withpartcontrol



withsoftcontrol bool yes yes

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Soft proton background control, "yes" to subtract the soft proton background image.

withswcxcontrol bool yes yes

Solar wind charge exchange background control, "yes" to subtract the swcx background image.

withmaskcontrol yes bool yes

For masking with an additional image.

mask.fit mask dataset yes

Mask image file name.

clobber boolean T/Fno yes

Clobber existing files?

5. Task cheese parameters:

prefixm string

Detector and exposure identifiers (eg. "1S001 2S002") for the MOS exposures (in the example MOS1 S001 and MOS2 S002) to be processed.

prefixp yes string Detector and exposure identifiers (eg. "S003") for the PN exposures (in the example PN

S003) to be processed.

verb yes int 4

SAS verbosity level.

scale yes real 0.5

Energy fraction, which sets the exclusion radius of point sources.

1.0 yes real

Flux threshold (in units of 1.0E - 14cgs for the exclusion of point sources.

dist yes real Minimum separation in arc seconds between masked sources.

elow 400 int

The low energy for the band in eV

int 1250

The high energy for the band in eV

clobber boolean T/Fno yes

Clobber existing files?

6. Task cheese_bands parameters:

prefixm	yes	string						
Detector and exposure ide	entifiers (e	g. "1S001	2S002")	for th	e MOS	exposures	(in the ex-	-
ample MOS1 S001 and M	OS2 S002)	to be prod	cessed.					

prefixp	yes	string	



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Detector and exposure identifiers (eg. "S003") for the PN exposures (in the example PN S003) to be processed.

verb	yes	int	4	
SAS verbosity level.				
scale	yes	real	0.5	
Energy fraction, which	h sets the ex	clusion radiu	s of point sources	•
ratet	no	real	1.0	
Total flux threshold for	or exclusion	of pt srcs		
rates	no	real	1.0	
Soft flux threshold for	exclusion o	f pt srcs		
rateh	no	real	1.0	
Hard flux threshold for	or exclusion	of pt srcs		
1.		,	T	
dist	no	real		
Minimum separation i	in arc second	ds between m	asked sources	
elowlist		1 . ,	400 2000	
	yes	int	400 2000	
Lower energy limit lis	t for the ene	ergy bands in	eV	
ehighlist	TOG	int	1300 7200	
	yes			
Higher energy limit lis	st for the en	ergy bands n	1 e v	
clobber	no	boolean	yes	T/F
CIODDCI				

- 7. Task clean parameters: none
- 8.

Task comb parameters:							
caldb	yes	string					
Directory containing all the ESAS specific calibration files							
	ı.	1					
withpartcontrol yes boolean true							
Particle background flag,	Particle background flag, 'true' to include it.						
withsoftcontrol	yes	boolean	true				
Soft proton background flag, 'true' to include it.							
withswcxcontrol	yes	boolean	true				
SWCX background flag 'true' to include it							

SWCX background flag,	'true' to include it.	

alpha	yes	real	1.7	
1 1 1 1 6	+1 C1+		1.	

Assumed spectral index for the filter correction scaling.

elowlist	yes	int	400 750	
Energy low limit(a) (in al.	() for the	ifferent he	nda	

Energy low limit(s) (in eV) for the different bands.

ehighlist	yes	int	750 1250	
		11.00		

Energy high limit(s) (in eV) for the different bands.



mask int 0 yes Masking control. 0: No additional masking, 1: uses the mask produced by the cheese task, 2: uses the normal mask images produced by eexpmap, and 3: uses the normal mask images produced by eexpmap modified by make-mask. 1S001 2S002 S003 prefixlist yes string "1S001 2S002 S003") for the exposures (in the example MOS1 Exposure identifiers (eg. S001, MOS2 2S002, and PN S003) to be processed. T/F clobber no boolean Clobber existing files? 9. Task conv_reg parameters: detector string yes The instrument identifier (EMOS1, EMOS2, or PN). mode yes intconv_reg operational mode: mode=1 - region fits files are both input and output mode=2 – ascii files with region parameters are both input and output mode=3 - command line input of individual region parameters and screen output imagefile yes string Filename image in sky coordinates – used to extract observation position angle. rano real none mode=3 RA input \mathbf{dec} no real none mode=3 Dec input real no mode=3 region shape input, only circle and ellipse at this time (either all upper or all lower case, along with their "nots", e.g., "!ELLIPSE") radius no real none mode=3 radius for circular region input semimajor no real none mode=3 semimajor axis (in arc minutes) for elliptical region input semiminor no real none mode=3 semiminor axis (in arc minutes) for elliptical region input rotangle none no real mode=3 rotation angle (in degrees) for elliptical region input inputfile no string mode=1,2 input file name

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mode=1,2 input file name

no

string

outputfile



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10. Task esas-mosaic parameters:

sublist	yes	string	mosaic-times.txt	
Ascii file with the output	of esprep_i	$mosaic_tim$	es	
m1prefix	yes	string	1S001	
MOS1 ESAS prefix			•	
m2prefix	yes	string	2S002	
MOS2 ESAS prefix	'	•	,	
pnprefix	yes	string	S003	
PN ESAS prefix				

11. Task espfilt parameters:

eventset	no	string		none
list of event files	l	_		
method	no	string		corner
which method to use.				
withsmoothing	no	boolean	N	Y/N
Smooth data?	110	Boolean	11	1/11
Sillottii data.				
smooth	no	integer	50	> 1
Smoothing factor in secon	ds			
				77/27
withbinning	no	boolean	N	Y/N
Bin data?				
binning	no	integer	50	> 1
Bin width in seconds	110	integer	00	/ 1
Din width in seconds				
withspecranges	no	boolean	N	Y/N
Use upper/lower spec cha	ns?			
specchanmin	no	integer	2500	> 1 ev, < 32766
Low Spectral Channel				
specchanmax	no	integer	12000	> 2 ev, < 32767
High Spectral Channel				
ratio	no	real	1.2	> 0.01, < 10.0
Flaring ratio of annulus_c	nts corn_a	rea corn_c	enst annu_area	
clobber	no	boolean	yes	T/F
Clobber existing files?	L	<u>I</u>	, v	,

12. Task esprep_mosaic parameters:

eventsin	yes	string	mos1S001-	
			clean_P001.ds	

Input event list from emosaic_prep.



Input mask file name.

yes

real

flimtot

mos1S001eventsout yes string clean.fits Output event list for use with ESAS. real 270.0 Reference RA of the sub-pointing. decobj 45.0 real Reference Dec of the sub-pointing. clobber no boolean T/F yes Clobber existing files? 13. Task esprep_mosaic parameters: atsfile string Input ODF *ATS.FIT file. sasfile string Input summary ODF *SUM.SAS file created by odfingest. 14. Task make_mask parameters: inimage string inimage.fit Event image for the exposure inmask inmask.fit string no Exposure mask outmask string outmask.fit no The output file name for the cheese mask reglist string reglist.fit The filtered source region list. clobber T/F no boolean yes Clobber existing files? 15. Task make_mask_merge parameters: srclistmerged-sourceyes string list.fits Merged source list from merge_source_list prefix 1S001 yes string Exposure identifier. inmask string mos1S001-maskyes im-750-1250.fits

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Combined band	source flux	threshold	(10^{-14})	cgs).	•
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		,	5 /	
flimsoft	yes	real		
Soft band source flux three	shold (10	color colo		
	,	- ,		
flimhard	yes	real		
Hard band source flux thr	eshold (10	-14 cgs).		
	`	,		
scale	yes	real		
Scale factor for W90 radiu	ıs.		,	
seper	yes	real		
Minimum allowed source s	separation	in arc seco	ond.	
maxlikelim	yes	real		
Minimum accepted value	for the ma	ximum like	elihood detetion paran	neter.
clobber	no	boolean	yes	T/F

Clobber existing files?

16. Task merge_comp_xmm parameters:

caldb	yes	string	

Directory containing all the ESAS specific calibration files

14 01			1.	
dirfile	yes	string	mydir	

File containing the list of exposures, including the directory string, for data to be merged. For instance, if the merging processing is being done in the directory /DATA/merge and the individual observations are located in the parallel directories /DATA/obs1/proc and /DATA/obs2/proc, the file dirlist could have entries such as:

/DATA/obs1/proc/mos1S001 /DATA/obs1/proc/mos2S002 /DATA/obs1/proc/pnS003 /DATA/obs2/proc/mos1S001 /DATA/obs2/proc/mos2S002 /DATA/obs2/proc/pnS003

coord	yes	int	1	
Sologta which goordinate	watom aho	uld be used	d 1. polintia 2. pount	orial 3. calactic

Selects which coordinate system should be used, 1: ecliptic, 2: equatorial, 3: galactic.

crvaln1	yes	real				
Control langitude of the projection						

Central longitude of the projection.

crvaln2	yes	real	
0 . 11 1			

Central latitude of the projection.

pixelsize yes	real		
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Pixel size of the projection in decimal degrees.

component	yes	int	1	

Component to be cast, 1: count image, 2: exposure, 3: QPB counts, 4: SP counts...



alpha real 1.7 yes Assumed spectral index for the filter correction scaling. 400 yes Energy low limit (in eV) for the band. ehigh 1250 yes Energy high limit (in eV) for the band. maskcontrol Mask control, 0: no masking, 1: point source masking using the output from cheese, 2: good area masking using the masks produced by mos-spectra, 3: mask from merged source list output from make_mask_merge. xdim 2000 intX dimension of the output image int 2000 Y dimension of the output image clobber T/F boolean no yes Clobber existing files?

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17. Task merge_source_list parameters:

File containing the list of ObsID directory strings for source lists to be merged. For instance, if the merging processing is being done in the directory /DATA/merge and the individual observations are located in the parallel directories /DATA/obs1/proc and /DATA/obs2/proc, the file *dirlist* could have entries such as:

/DATA/obs1/proc/ /DATA/obs2/proc/

maxlikelim	yes	real			
Minimum accepted value for the maximum likelihood detetion parameter					

clobber	no	boolean	yes	T/F

Clobber existing files?

18. Task mos_back parameters:

prefix	yes	string	S003			
D / / 1 '1 / 'C / 10001 C MO01 0001						

Detector	and	cxposure	identifica	ı,(c.g.,	10001	101	MODI	DOOL	cxposurc)	•

caldb		yes	string			
D: 4 11.41 TOA O						

Directory containing all the ESAS specific calibration files

diag	yes	$_{ m int}$	1	
Controls the amount of di	omnostia o	utnut (0.10	w 1 modium 2 high)	

Controls the amount of diagnostic output (0 low, 1 medium, 2 high)

elow	yes	int	400	
Energy low limit (in eV) f	or the ban	d.		



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ehigh	yes	int	1250								
Energy high limit (in eV) for the band.											
$\operatorname{ccd}[1-7]$	yes	int	1								
Selects which ccd's should	Selects which ccd's should be included.										
clobber	no	boolean	ves	T/F							

Clobber existing files?

- 19. Task mos-filter parameters: None
- 20. Task mos-spectra parameters:

prefix	yes	string	1S001	
Detector and exposure ide	entifier (eg.	"1S001")	for MOS1 S001 e	exposure to be processed.

caldb	yes	string	

Directory containing all the ESAS specific calibration files

region	yes	\inf	reg.txt	

the selection expression for the desired region for the generation of the model background spectrum. If no file with the input name exists, or if the file is empty, then the default is to model the data from the entire field of view. If a specific region is desired, the region expression must be in detector coordinates. For example, a file containing &&((DETX,DETY) IN circle(201,-219,3600)) would extract the central 3' of the cluster Abell 1795. Note that the leading "&&" are required as the selection expression is added to other constraints.

mask		yes	int	0					
				1 .	1 .1	.11			

Flag to mask out point sources. O selects no masking while 1 will cause mos-spectra to use the output filtered source region file from cheese or cheese-bands.

elow	yes	int	400	

Energy low limit (in eV) for the band. If elow and ehigh are set to 0, the image processing will be eliminated and only spectral files will be produced.

ehigh	yes	int	1250	

Energy high limit (in eV) for the band. If *elow* and **ehigh** are set to 0, the image processing will be eliminated and only spectral files will be produced.

ccd1-7 yes	int	1	
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Flag to include individual CCDs. 1 to include, 0 to not.

21. Task pn_back parameters:

prefix	yes	string	S003	
Detector and exposure ide	entifier.(e.g	S003 ext	oosure).	

caldb	yes	string		

Directory containing all the ESAS specific calibration files

diag	yes	int	1	
Controls the amount of di	agnostic or	utput (0 lo	w, 1 medium, 2 high)	



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elow	yes	int	400		
Energy low limit (in eV) for the band.					
ehigh	yes	int	1250		
Energy high limit (in eV) for the band.					
quad[1-4]	yes	int	1		
Selects which PN quadrar	ts should l	e included	1		

Selects which PN quadrants should be included.

clobber no boolear	yes	T/F
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Clobber existing files?

22. Task pn-filter parameters: None

23. Task pn-spectra parameters:

prefix	yes	string	1S001	
Detector and exposure ide	entifier (eg.	"S001") f	for the PN S001 expos	ure to be processed.

caldb	yes	string		

Directory containing all the ESAS specific calibration files

region	yes	int	reg.txt	
11 1 1 · · · · · · · · · · · · · · · ·	/1 1 ·	1 .	C 11 1 C	1 111 1 1

the selection expression for the desired region for the generation of the model background spectrum. If no file with the input name exists, or if the file is empty, then the default is to model the data from the entire field of view. If a specific region is desired, the region expression must be in detector coordinates. For example, a file containing &&((DETX,DETY) IN circle(201,-219,3600)) would extract the central 3' of the cluster Abell 1795. Note that the leading "&&" are required as the selection expression is added to other constraints.

mask	yes	int	0	

Flag to mask out point sources. θ selects no masking while θ will cause pn-spectra to use the output filtered source region file from cheese-bands.

elow	yes	int	2000	

Energy low limit (in eV) for the band. If *elow* and *ehigh* are set to θ , the image processing will be eliminated and only spectral files will be produced.

ehigh	yes	int	7200	

Energy high limit (in eV) for the band. If *elow* and *ehigh* are set to θ , the image processing will be eliminated and only spectral files will be produced.

quad1-4	yes	int	1	

Flag to include individual quadrants.

24. Task point_source parameters:

runc yes string mateos	func yes		
------------------------	----------	--	--

Name of logN-logS function.

hms (Hasinger, Miyaji, & Schmidt 2005)

mushotzky (REF TBD)

cappelluti (Cappelluti et al. 2008)



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mateos (Mateos et al. 2008)

yes	real	1.0e-14	
	Tear	1.00-14	
/ CIII2/ 5.			
yes	real	10.6	
smic X-ray	background		
ves	real	1.46	
Σ.			I
:			
yes	string		
identifiers ((eg. "1S001	") for the MOS exp	osure S001) to be p
VOS	string		
		hration files	
	specific can	oration mes	
yes	string	1	
ot (0) a CC	D.		
. ,			
yes	int	400	
band in eV	'		
yes	int	1250	
band in eV	7		
VAS	int	1	
	-	-	
., 2 101 @ 510	oken power	TOW	
no		0	
only if spec	ctrumcontro	l=1	
no		0	
w index, or	aly if spectr	umcontrol=1	
no		0	
	only if spe	-	
i iaw ilidex,	, omy ii spe	cti umcontroi—2	
no		0	
	model, only	"	<u> </u>
	model, only	$\begin{array}{ c c c c c c c c c c c c c c c c c c c$	-2
	model, only	"	-2 -2
no no		y if spectrumcontrol=	-2 -2
no er law index		y if spectrumcontrol= 0 ectrumcontrol=2	-2
no er law index	x, only if sp	y if spectrumcontrol= 0 ectrumcontrol=2	=2
no er law index	x, only if sp	y if spectrumcontrol= 0 ectrumcontrol=2	-2
	yes smic X-ray yes smic X-ray yes yes identifiers yes the ESAS s yes ot (0) a CC yes band in eV yes band in eV no only if spec no only if spec	yes real smic X-ray background yes real : yes string identifiers (eg. "1S001 yes string the ESAS specific cali yes string ot (0) a CCD. yes int band in eV yes int band in eV yes int cali yes string ot (0) a CCD.	yes real 10.6 smic X-ray background. yes real 1.46 : yes string dentifiers (eg. "1S001") for the MOS exp yes string the ESAS specific calibration files yes string 1 ot (0) a CCD. yes int 400 band in eV yes int 1250 band in eV yes int 1 , 2 for a broken power law no 0 only if spectrumcontrol=1

Clobber existing files?



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26.	Task	proton_scale	parameters:

caldb yes string	
------------------	--

Directory containing all the ESAS specific calibration files

mode	yes	int	1	

mode - 1: do a single region, 2: do multiple regions with the required input provided in a text file (parameter spfile).

det yes int	1	1—2—3
-------------	---	-------

FOR MODE=1 - Detector, 1 for MOS1, 2 for MOS2, and 3 for PN

maskfile	yes	string	region	

FOR MODE=1 - File name for the mask file. This is the mosprefix-obj-im-sp-det.fits file produced for the region by mos-spectra.

	specfile	yes	string	specfile	
--	----------	-----	--------	----------	--

FOR MODE=1 - File name for the spectral file for the region.

spfile yes string

FOR MODE=2 - ASCII text file with the input for multiple regions. The file should contain, on separate lines, the detector number (det), mask file name (mask), and spectral file name (spec) for each region.

27. Task rot_det_sky parameters:

mode	yes	int	1	
Selection on particle (1), s	oft proton	(2). SWCX	(3) backgrounds. (4)	MASK. (5) MASKIT.

prefix	yes	string	1S001		
D / 1	· 1 / · C /	" 1 COO1	22 \ C +1	MOO	0001) / 1

Detector and exposure identifiers (eg. "1S001") for the MOS exposure S001) to be processed.

	elow	yes	int	350	
--	------	-----	-----	-----	--

The low energy for the band in eV

ehigh	yes	int	800	

The high energy for the band in eV

detx yes int 0	detx
----------------	------

The DETX reference pixel location

dety	yes	int	0	

The DETY reference pixel location

	$\mathbf{s}\mathbf{k}$	$\mathbf{y}\mathbf{x}$		yes	$_{ m int}$	0	
- 7	T 1	37 1	 C . 1				

The X location of the reference pixel

skyy yes int 0	
----------------	--

The Y location of the reference pixel location

m	askfile			yes	string									
	0.1	c	•		 1 11	1	1 .	1	•	- 1	TC 1 C: 1 1	1 . 1	. 1	

The file name for an image to provide additional masking if desired. If left blank then there will be no additional masking. The mask images must be the same size and projection of the other images.



clobber T/F no boolean yes Clobber existing files? 28. Task rotimdetsky parameters: prefix string 1S001 ves "1S001") for the MOS exposure S001) to be pro-Detector and exposure identifiers (eg. cessed. mask string none yes The file name for an image to provide additional masking if desired. If left blank then there will be no additional masking. The mask images must be the same size and projection of the other images. elow 400 int The low energy for the band in eV 1250 ehigh int The high energy for the band in eV mode yes int Selection on particle (1), soft proton (2), SWCX (3) backgrounds, (4) MASK, (5) MASKIT. clobber T/F no boolean yes Clobber existing files? 29. Task sp_partial parameters: caldb 1S001 yes string Directory containing the ESAS calibration files. detector yes int Detector to be processed 1-MOS1, 2-MOS2, and 3-PN. mos1S001-spfullimage yes string ps.fits Image from the full field of view. fullspec mos1S001-objyes string ps.pi Spectrum from the full field of view. regionimage mos1S001-spyes string nps.fits Image from the selected region. regionspec yes string mos1S001-objnps.pi

0.05

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Page:

Xspec normalization of the SP component.

ves

real

Spectrum from the selected region.

rnorm

30. Task swcx parameters:

1	1									
prefix	yes	string								
Detector and exposure identifiers (eg. "1S001") for the MOS exposure S001) to be pro										
cessed.										
caldb	yes	string								
Directory containing all t	he ESAS s		bration files							
ccd[1-7]	yes	string	1							
Flag to include (1) or not			1							
r lag to merade (1) or not	(0) a CC1	J.								
elow	7700	int	400							
I	yes	1110	400							
The low energy for the band in eV										
1 • 1			1200							
ehigh	yes	int	1300							
The high energy for the b	oand in eV									
elinelist	yes		1 2							
Energies of SWCX lines t	o be includ	ded								
gnormlist	yes		0.1 0.03							
Gaussian normalizations	from Xspec	3								
objrmf	yes	string								
RMF for the region										
<u> </u>										
objarf	yes	string								
ARF for the region	J	8								
objspec	yes	string								
Spectrum for the region	J 05	2011118								
Spectrum for the region										
clobber	no	boolean	1705	T/F						
	110	Doolean	yes	1/1						
Clobber existing files?										

5 Errors

This section documents warnings and errors generated by this task (if any). Note that warnings and errors can also be generated in the SAS infrastructure libraries, in which case they would not be documented here. Refer to the index of all errors and warnings available in the HTML version of the SAS documentation.

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6 Input Files

- 1. adapt: maskfile (Input file for additional masking).
- 2. adapt_merge: maskfile (Input file for additional masking).
- 3. bin_image: maskfile (Mask file name).
- 4. make_mask: inimage (The event image for the exposure).
- 5. make_mask: inmask (Exposure mask).
- 6. make_mask: reglist (The filtered source region list).
- 7. mos-spectra: region (File with additional region information).
- 8. proton_scale: region (Mask image file name).
- 9. proton_scale: specfile (Spectrum file name).
- 10. proton_scale: spfile (ASCII text file with the input for multiple regions).
- 11. rot_Det_Sky: maskfile (Input file for additional masking).
- 12. sp_Partial: fullimage (Full region image file name).
- 13. sp_Partial: fullspec (Full region spectrum file name).
- 14. sp_Partial: regionimage (Selected region image file name).
- 15. sp_Partial: regionspec (Selected region spectrum file name).

7 Output Files

See individual routine documents in this directory for output file descriptions.

8 Algorithm

9 Comments

We would like to thank members of the MOS and pn hardware and software teams, the XMM-Newton SOC at the European Space Astronomy Center (ESAC), and other members of the EPIC Background Working Group for their contributions which ranged from helping us to understand instrument and software issues to the identification of filter-wheel closed observations in the archive. Users of this package should be aware of the informational web pages covering the background issues of EPIC observations at: http://xmm.esac.esa.int/external/xmm_sw_cal/background/index.shtml and the EPIC Calibration Status document at:

http://xmm.esac.esa.int/external/xmm_sw_cal/calib/index.shtml.



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- [3] Snowden, S. L., Kuntz, K. D., Cookbook for Analysis Procedures for XMM-Newton EPIC MOS Observations of Extended Objects and the Diffuse Background, 2008, Vol 3.0.

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