

## esplinemap

February 1, 2016

#### Abstract

ESPLINEMAP uses the **eboxdetect** local mode source list to derive spline background maps from the non-source regions.

### 1 Instruments/Modes

Instrument	Mode
EPIC MOS	IMAGING
EPIC PN	IMAGING

#### 2 Use

pipeline processing	yes	
interactive analysis	yes	

### 3 Description

Sources found in the local detection step (task **eboxdetect**) at significance levels (column SIGMA of eboxdetect source list) exceeding a user-specifiable threshold (input parameter mlmin) in the respective energy band are removed from the image using a suitable PSF and source brightness dependent cutout radius (determined to be the radius at which each source contributes more than a user-specifiable number of counts/arcsec<sup>2</sup> to the background; parameter scut; default value: 0.01). The resulting image can optionally be output for diagnostic purposes. After the removal of the sources the image is rebinned to a grid of the dimension nsplinenodes  $\times$  nsplinenodes.

Division of the image by the exposure image removes gradients due to spatial variations of the exposure from the image which otherwise would not be well represented by the spline fit. A two dimensional spline fit of the rebinned and exposure corrected image is performed. The number of spline nodes (default value: 16) is user-selectable. Finally, the resulting spline image is again multiplied by the exposure images. If the parameter **nfitrun** is set to values > 1, remaining excesses of the input image over the result of spline fit can be removed iteratively: if pixels of the rebinned image deviate from the spline fit by more than a specifiable number of sigmas (default value: 4 sigmas) the excesses are removed by setting their statistical weights to zero and the spline fit is repeated (maximum number of iterations may be specified). The number of removed bins and the reduced chi<sup>2</sup> values are displayed when using verbosity level 5 or



higher. Note that removal of a large number of contiguous bins will lead to areas where the spline fit is unconstrained.

The reduced chi<sup>2</sup> and corresponding number of degrees of freedom of the background map with respect to the input image is stored in the keywords CHISQR and NDOF of the output background map.

From version 3.0 esplinemap is able to determine the background caused by out-of-time events registered during the readout process of the PN CCDs. If the flag withootset is set, the photon event table specified in **ooteventset** is read and the background caused by OOT events is included in the output background map. As input table **esplinemap** can use either a normal photon event data-set or a photon events table created with **epevents** with flag withoutoftime set. Note that in both cases a photon event set has to be filtered with the same temporal and flag selections as the image used as input to **esplinemap**. The parameters **pimin** and **pimax** are used to specify the energy range of the input image and to select those photons from the input event list that fall into this energy range. If the input event table contains only photons within the energy range of the input image, the parameters can be left at their default values **pimin=1** and **pimax=30000**.

With version 4.0 an alternative method to fit the background of an image has been implemented: If the option fitmethod=model is set, a 2-component model for the detector (particle) and the cosmic X-ray backgrounds is fit to the masked and binned input image. This model consist of a linear combination of the vignetted exposure map and the unvignetted exposure mask of the input image. The exposure maps are specified by the user via the parameters expimageset and expimageset2. An example call of eexpmap and esplinemap is given here:

```
eexpmap imageset=image.fits eventset=events.fits attitudeset=attitude.fits \
    withvignetting=yes expimageset=expmap1.fits pimin=500 pimax=2000
eexpmap imageset=image.fits eventset=events.fits attitudeset=attitude.fits \
    withvignetting=no expimageset=expmap2.fits pimin=500 pimax=2000
esplinemap imageset=image.fits boxlistset=eboxlist.fits withexpimage=yes \
    bkgimageset=bkg_model.fits \
    withexpimage2=yes expimageset=expmap1.fits expimageset2=expmap2.fits \
    pimin=500 pimax=2000 \
    fitmethod=model
```

If only one exposure map is provided, it is assumed to be a vignetted exposure map and a flat image is used as the second model component. All other parameters of **esplinemap** can be used as in the case of fitmethod=spline.

#### 4 Parameters

This section documents the parameters recognized by this task (if any).

Parameter	Mand	Type	Default	Constraints
1			1	
boxlistset	yes	filename	eboxlist.fits	
Name of eboxdetect source list				
imageset	yes	filename	image.fits	
Names of input EPIC fits images				



bkgimageset	yes	filename	bkgimage.fits	
Name of output spline bac	kground map	)		· ·
expimageset	no	filename	expimage.fits	
Name of input exposure m	ар			
expimageset2	no	filename	expimage.fits	
Name of second input expo	-			
1 1	1	Ŧ		
detmaskset	no	filename	detmask.fits	
Name of input detection m	ıask			
scut	no	float	0.01	$[0.0 \le param \le 10.0]$
Source cut-out flux level; [	-		0.01	
	sounds, areso	~ ]		
idband	no	integer	1	$[0 \le param \le 9]$
Band id from eboxdetect s	ource list	·		
1.			1.0	
<b>mlmin</b> Minimum single band dete	no no	float	1.0	$[0.0 \le param \le 50.0]$
Minimum single band dete	ction intenno	ou loi source	s to be cut out	
nsplinenodes	no	integer	16	$[10 \le param \le 20]$
Number of spline nodes	I			
-				
excesssigma	no	float	4.0	$[1.0 \le param \le 6.0]$
Threshold for excesses sign	nas with resp	ect to backgr	ound spline fit	
nfitrun	no	integer	3	$[1 \le param \le 5]$
Number of iterations for re		0		$[1 \leq param \leq 0]$
		obbob, infor an		
withdetmask	no	boolean	true	
Flag to use detection mask		L		
withexpimage	no	boolean	true	
Flag to use exposure map	110	boolean	truc	
1 108 to use emposare map				
withexpimage2	no	boolean	true	
Flag to use second exposur	e map for op	otion "fitmeth	od=model"	
with aboard		haalaan	false	
withcheese	no f photon imp	boolean		out (socalled cheesed image)
Controls optional output o	i photon ina	ige where soul	ices have been masked	out (socalied cheesed image)
cheeseimageset	no	filename	cheese.fits	
Name of diagnostic output	cheesed ima	ge.		
·				
withcheesemask	no	boolean	false	
Controls optional output o	t mask image	e, value=0 for	masked areas, value=	1 for valid image areas
cheesemaskset	no	filename	cheesemask.fits	
Name of diagnostic output	-		011005011105K.11105	



ooteventset	no	filename	events.fits		
Name of photon event table used for background due to out-time-events					
pimin	no	integer	1	[1 < param < 30000]	
Lower PI channel limits of input images.					
pimax	no	integer	30000	[1 <param<30000]< td=""></param<30000]<>	
1		0		[ [	
Upper PI channel limits of inj	put images.	0		[- (Fortonic (00000)]	
-	put images.			[- (Farmer (00000)]	

Fitting method: spline fit or 2-component background model.

#### 5 Errors

This section documents warnings and errors generated by this task (if any). Note that warnings and errors can also be generated in the SAS infrastructure libraries, in which case they would not be documented here. Refer to the index of all errors and warnings available in the HTML version of the SAS documentation.

SplineFitError (error) Less than 1 data point for chi\*\*2

- SplineFitError (error) Error condition in spline fit
- SplineFitError (error) Error in LPSROS
- SplineFitError (error) Error in CALBCS
- SplineFitError (error) Error in ERPLPS

WrongType (error) Input image has wrong type

- WrongRefPix (error) Reference pixel is outside of FOV
- FileMismatch (error) Exposure map has wrong instrument
- **FileMismatch** (error) OOT list is not an EPN event list

FileMismatch (error) Exposure map has wrong size

FileMismatch (error) Detection mask has wrong size



#### 6 Input Files

- 1. PPS product (from task **eboxdetect** run in local detect mode): EPIC eboxdetect local mode source list
- 2. PPS product (from task evselect): EPIC FITS image
- 3. PPS product (from task **eexpmap**): EPIC exposure image
- 4. from task emask: Detection mask
- 5. Filtered photon event list from **evselect**.

### 7 Output Files

- 1. Spline background map (to be used by the tasks **eboxdetect** (map mode), **emldetect**, and **esensmap**)
- 2. Photon image where sources have been removed (so-called cheesed image; optional diagnostic output)

### 8 Algorithm

subroutine esplinemap

```
    If out-of-time events flag is set: Calculate background due to
OOT events and subtract from image.
    Remove source count dependent circular area around EBOXDETECT
```

```
(local mode) sources
which have detection likelihoods greater than a specifiable
threshold from the image.
```

- 3) If exposure flag is set, divide image by exposure map.
- Perform spline fit of image, using a user-specifiable number of spline nodes. If detection mask flag is set, only use image regions marked by detection mask.
- 5) If exposure flag is set, multiply spline image by exposure map.

```
6) IF rebinned image pixels contain excesses above spline fit
THEN
remove excesses from image and repeat (2) - (5) a specifiable
number of times
END IF
7) If out-of-time events flag is set: Add OOT events background to background map.
```

end subroutine esplinemap



# 9 Comments

10 Future developments

References