



# eupper

February 1, 2016

## Abstract

A routine to calculate the count rate upper limit at a position within an image

## 1 Instruments/Modes

Instrument	Mode
EPIC	IMAGING

## 2 Use

pipeline processing	no
interactive analysis	yes

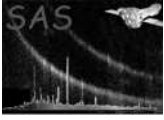
## 3 Description

This task is designed to analyse a circular region within an image, specified in celestial coordinates, and calculate the upper limit at that position. It has been designed to run as quickly as possible and to that end avoids using any calibration quantities. The input image should be in sky coordinates and WCS keywords in the header are used to find the required position within the image. The task should run on any sky image regardless of the instrument.

A more flexible but slower result for XMM-Newton EPIC images, using the full XMM-Newton calibration, can be obtained from the **eregionanalyse** task.

### 3.1 Encircled energy function

The point spread function (PSF) is a function of off-axis angle and photon energy for all X-ray cameras. The correct calculation of this value is time consuming. In practise, the XMM-Newton EPIC cameras have been designed to minimise the variation of the PSF, and hence the encircled energy (EE), across the field of view and across the range of observed photon energies. For this reason, a good approximation to the EE, for fixed radii, can be tabulated within the code for each camera. The tabulated values, used within the program, are shown below but may be overridden on the command line. They should be accurate to about 5%.



Extraction radius (arcsec)	Camera			
	MOS-1	MOS2	EPN pointed	EPN slew
15	0.68	0.69	0.71	0.65
30	0.83	0.83	0.88	0.85
45	0.89	0.89	0.93	0.91
60	0.92	0.93	0.95	0.94

## 3.2 Input

### 3.2.1 Source image

The source image should be input in sky (X/Y) or celestial (RA/DEC) coordinates. The task expects to find astrometry keywords in a certain part of the primary header and will exit with an error if the keywords are not found. Standard images produced by `evselect` and `xmmselect` and the pipeline will process ok. Specifically images need to have the coordinate transformation (CTYPE1 keyword) as "RA—TAN" or "EQU—CAR".

### 3.2.2 Source region

The upper limit will be found from a circular region within the input image, centred on the celestial (RA, DEC, FK4 2000) position defined by the parameters `srcra` and `srcdec` with radius given by `srcradius`.

### 3.2.3 Background

Background is taken from an annular region within the source image via the `bgdra`, `bgdra`, `bgdrinner` and `bgdrouter` parameters.

### 3.2.4 Exposure

By default the exposure time used for converting counts to count rate is read from the `EXPOSURE` keyword in the header of the source image. An exposure map may be introduced by the `exposuremap` parameter in which case the exposure time is taken as the value at the central pixel of the source region within the exposure map.

## 3.3 Output

### 3.3.1 Statistical upper limit

By default this gives the 2-sigma upper limit on the background subtracted source count rate. If the number of counts is small, i.e. there are less than 80 counts in the source region then the Bayesian upper limit is returned (see Kraft, Burrows and Nousek, 1991, ApJ 374, 344). For larger numbers of counts the returned upper limit is given by:

$$U = (\max((S - B), 0) + 2 * \sqrt{S + B * (A_S/B_S)^2}) / \text{exp\_tim} \quad (1)$$



Note this number is the upper limit for the count rate in the input source region before correction for encircled energy.

The significance level of the upper limit may be changed using the `ulsig` parameter.

### 3.3.2 Other details

If the parameter `details` is true then the following information is also output:

Source/back counts: number of photons in the source and backgnd regions

Area ratio: the ratio of the source to background region areas

Exposure: The exposure time at the centre of the source region

EEF: The encircled energy factor used in the calculation

## 3.4 Output format

The output format shown below is dependent on the value of the `details` keyword. The strings shown may be searched for in a script and every effort will be made to keep them constant between versions of this task.

```
details = true
```

```
eupper:- Source/back counts: 1554 4599
Area ratio: 0.253
Exposure: 11980.7
EEF:0.95
eupper:- Statistical upper limit c/r: 0.041501066 c/s
```

```
details = false
```

```
eupper:- Statistical upper limit c/r: 0.041501066 c/s
```

## 3.5 Examples

### 3.5.1 Find an upper limit using an annular background region

```
eupper imageset=imagepnew.ds srcra=187.912 srcdec=64.2 srcradius=0.00833 bgdra=187.912 bgddec=64.2
bgdrinner=0.01666 bgdrouter=0.03333
```

This will find the upper limit from a 30 arcsecond circle about the sky position, `ra=187.912 degs`, `dec=64.2 degs`, using a background annulus about the source circle of inner radius, 1 and outer radius 2 arcminutes. The exposure time will be taken from the primary header of the file `imagepnew.ds` and the result will be a 2-sigma upper limit.



### 3.5.2 Use an exposure map

```
eupper imageset=emosimage.ds exposuremap=emosexpos.ds srcra=187.912 srcdec=64.2
```

Takes the exposure value from the pixel in the file emosexpos.ds which contains the position ra=187.912 degs, dec=64.2 degs.

### 3.5.3 Find a 3-sigma upper limit

```
eupper imageset=emosimage.ds ulsig=0.9973
```

By default the task finds the 2-sigma (95.4% confidence upper limit. Here ulsig is set to 0.9973, the 3-sigma confidence value.

### 3.5.4 Use own value of the encircled energy factor

```
eupper imageset=emosimage.ds eef=0.9
```

Tells the program that the source region contains 90% of the counts.

## 4 Parameters

This section documents the parameters recognized by this task (if any).

Parameter	Mand	Type	Default	Constraints
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<b>imageset</b>	yes	string	image	
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The name of the input image.

<b>exposuremap</b>	no	string	NotSet	
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The name of the exposure map

<b>srcra</b>	yes	double		
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The right ascension of the upper limit position in degrees.

<b>srcdec</b>	yes	double		
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The declination of the upper limit position in degrees.

<b>srcradius</b>	yes	double		
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The radius of the source area (degrees)

<b>bgdra</b>	yes	double		
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The right ascension of the background area in degrees.

<b>bgddec</b>	yes	double		
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The declination of the background area in degrees.

<b>bgdrinner</b>	no	double	0.0	
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The inner radius of the background area (degrees). Deafulted to 0 in the case of a circular region



<b>bgdrouter</b>	yes	double		
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The outer radius of the background area (degrees).

<b>ulsig</b>	no	double	0.954	
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The significance for the upper limit calculation. The default is two-sigma (0.954). Set the value to 0.68 for one sigma or 0.997 for three sigma etc.

<b>witheef</b>	no	boolean	no	
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Whether the encircled energy factor has been specified as a parameter

<b>eef</b>	no	double	1.0	
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If `witheef=true` then the encircled energy factor is set by this parameter.

<b>details</b>	yes	boolean	no	
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Output details of the calculation including number of counts etc.

<b>withoutputfile</b>	yes	boolean	no	
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Write the output result to a text file ?

<b>output</b>	no	string	output.txt	
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Name of output text file

## 5 Errors

This section documents warnings and errors generated by this task (if any). Note that warnings and errors can also be generated in the SAS infrastructure libraries, in which case they would not be documented here. Refer to the index of all errors and warnings available in the HTML version of the SAS documentation.

### **DataNotSupported** (*error*)

The instrument is not recognised.

### **invalidArraySize** (*error*)

The image is not two dimensional.

### **InvalidWCSType** (*error*)

The image axes specified in the CTYPE1 and CTYPE2 keywords are not compatible.

### **InvalidPosWCSInfo** (*warning*)

The REFERENCE keywords in the image header, e.g. REF\_XCRPX, REF\_XCRVL, REF\_XCDLT are incomplete. Defaults are chosen but there is likely to be a problem later.

*corrective action:*

## 6 Input Files

- a sky coordinate image containing WCS keywords allowing a translation between celestial coordinates and image pixel. Coordinate systems RA—TAN and EQU—CAR are currently supported.



- optionally an exposure map containing WCS keywords allowing a translation between celestial coordinates and image pixel..

## **7 Output Files**

## **8 Algorithm**

## **9 Comments**

## **References**